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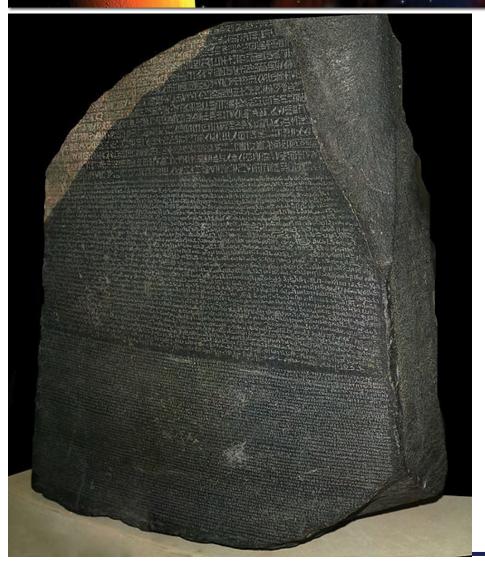
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Why is the Mission called Rosetta?



In 1799 a stone was found in the Egyptian city of Rosetta featuring writing in 3 languages from 196 BC. The text was written in hieroglyphs as well as ancient Greek. It took 23 years to translate the text.

Just like the Rosetta stone allowed us to understand Egyptian hieroglyphics the Rosetta spacecraft's visit to a comet will help us understand the origins of the solar system and the planet Earth.



Why is NASA Interested in the Study of Comets?

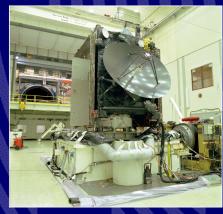
- Comets are preserved icy planetesimals, left over from the formation of the solar system
 - They contain a <u>chemical record</u> of the <u>conditions</u> in the solar nebula at the time of the formation of the planets (what was its temperature?)
 - There is strong evidence that comets <u>contain</u> unprocessed <u>material from</u> the natal interstellar cloud out of which the Sun and planets formed
- Comets maybe provided the volatile inventories of the terrestrial planets including pre-biotic molecules essential to the origin of life
 - The Late Heavy Bombardment was most likely the result of the clearing of comets from the outer planets zone
 - This bombardment provided a volatile veneer on the terrestrial planets including water and complex organics
- Long-period comets make up the more unpredictable fraction of the impact hazard at the Earth
 - An understanding of cometary structure and material strengths can only be achieved through rendezvous and lander missions

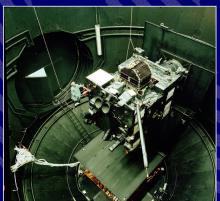




The Rosetta Spacecraft

The Spacecraft, Vital statistics





Size:

main structure 2.8x2.1x2.0 metres diameter of solar arrays 32 metres

Launch mass - total: 3000 kg (approx.)
- propellant 1670 kg (approx.)

- science payload 165 kg - Lander 100 kg

Solar array output 850 W at 3.4 AU, 395 W at

5.25 AU

Propulsion subsystem 24 bi-propellant 10N thrusters

Operational mission 12 years







Rosetta is designed to operate on solar power farther than any other spacecraft



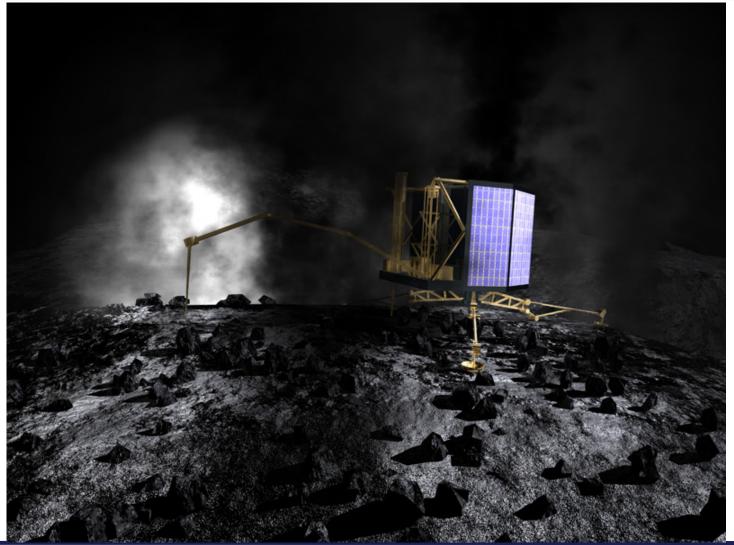


68 m2, 32m long, 22750 cells, from 8700 W close to Earth down to 440 W far away 5



Annual IAA Low-Cost Planetary Missions Conference

November 2014 Comet Landing







Rosetta Science Payload

Orbiter Teams

- 1. *ALICE UV spectrometer
- 2. CONSERT tomography/radio sounding
- 🛾 3. COSIMA chemistry 🧮
- 4. GIADA dust analysis
- 5. *IES ion and electron sensor
- 6. IPA plasma analyzer
- 7. MAP magnetometer
- 8. MIDAS atomic force
- 9. MIP magnetic impedance probe



- 10. *MIRO microwave spectrometer/radiometer
- 11. LAP Langmuir probe
- 🗆 12. OSIRIS camera 🌉
 - 13. *ROSINA mass spectromete
- 14. RSI radio science
 - 15. VIRTIS IR spectrometer

Lander Teams

- 16. APXS X-ray spectrometer, similar to that of Mars Pathfinder
- 17. CIVA -lander visible IR camera (omnidirectional)
- □ 18. COSAC lander mass spectrometer
- 19. MODULUS gas analyzer
- 20. MUPUS probe
- 21. ROLIS lander descent camera
 - 22. ROMAP lander magnetometer/material magnetism
 - 23. SESAME seismic data
- 24. CONSERT (2) tomography/radio sounding

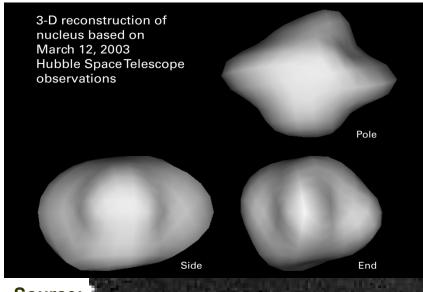
Legend:

US hardware contribution US investigation contribution



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The Target: Comet Churyumov-Gerasimenko



Source: CNES



Characteristics:

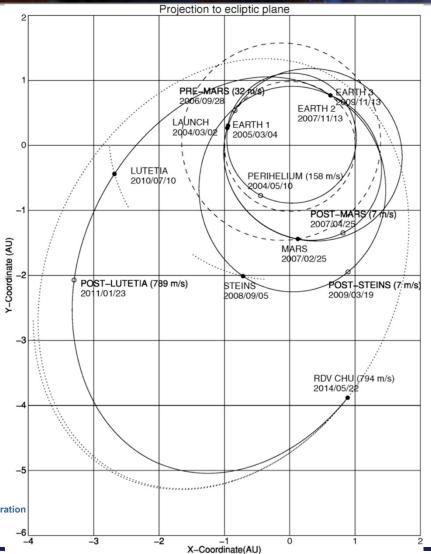
- Radius: 2.0 km

Rotation: 12.5 h

- unknown topography and surface properties
- temperatures
 - day ~ -50 ° C
 - night \sim -150 $^{\circ}$ C
- solar energy <1/10 of that near Earth
- gravity <10⁻⁵ g



Rosetta Trajectory





Oth Annual IAA Low-Cost Planetary Missions Conference

Science Targets to date

<u>Venus</u> MIRO April 1, 2004

Earth and Moon

Comet LINEAR

Comet Catalina

MAG and LAP.....June 26, 2005

Comet Tempel 1

Mars

Alice and IES Feb 25 – March 4, 2007

<u>Jupiter</u>

Alice.....Feb 27 – May 2007

<u>Asteroid</u> 2867/Steins Sept 3 – 2008

Asteroid 21/Lutetia July 10, 2011

National Aeronautics and Space Administration Jet Propulsion Laboratory California Institute of Technology

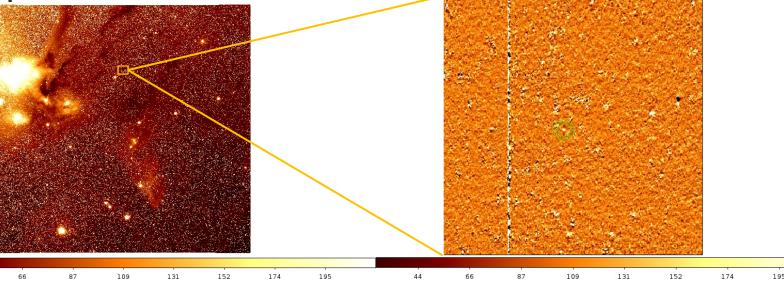


C-G image from Rosetta

This is the most recent picture from Rosetta. Both NAVCAMs were activated to image the comet and the sky that will be seen during the approach phase. The spacecraft downlinked the images acquired by the two cameras: OSIRIS and the NAVCAMs in preparation for the



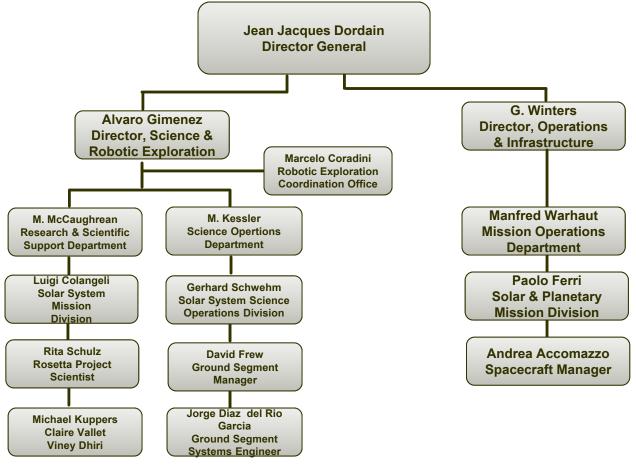
Source: ESA







ESA Reporting Chain for Rosetta







(C) In Annual IAA Low-Cost Planetary Missions Conference

US Rosetta Project Organization

Claudia Alexander Project Scientist

Art B. Chmielewski US Project Manager

Acquisitions

Joe Espinoza – Univ. of Maryland Richard Flores – SwRI Jane Lee - Lockheed Lisa Pham – University Contracts Krandalyn Richardson – WSU Susan Scrivner – Univ. of Arizona

Project Staff

Anna Marie Aguinaldo – Systems Engineer Andrea Angrum - Outreach Tracy Feehan – Resource Analyst Padma Varanasi– DSN Jose Macias – Mission Assurance Mgr. Monica Martin – Staff Assistant

ALICE Instrument Team

Alan Stern, PI – SwRI Joel Parker, PM – SwRI Mike A'Hearn, Co-I (Univ of Maryland) Hal Weaver, Co-I (John Hopkins) Paul Feldman, Co-I (John Hopkins)

IES Instrument Team

James Burch, PI (SwRI)
Tom Cravens, Co-I (Univ of Kansas)
Ray Goldstein, Co-I (SwRI)

MIRO Instrument Team Sam Gulkis , PI – JPL

Mark Allen, Co-I (JPL)
Margaret Frerking, Co-I (JPL)
Mark Hofstadter, Co-I (JPL)
Mike Janssen, Co-I (JPL)
Duane Muhleman, Co-I (MIT)
Peter Schloerb, Co-I (Caltech)
Tom Spilker, Co-I (JPL)

ROSINA

Stephen Fuselier Co-I (Lockheed Martin)
Tamas Gombosi Co-I (Univ. of Michigan)
Hunter Waite Co-I (SwRI)
Arthur Ghilmetti, Co-I (Lockheed Martin)
Dave Young, Co-I (SwRI)
Len Fisk, Co-I (Univ. of Michigan)
George Carrigan, Co-I (Univ. of Michigan)
Ed Shelley, Co-I (Lockheed Martin)

IDS Paul Weissman (JPL)

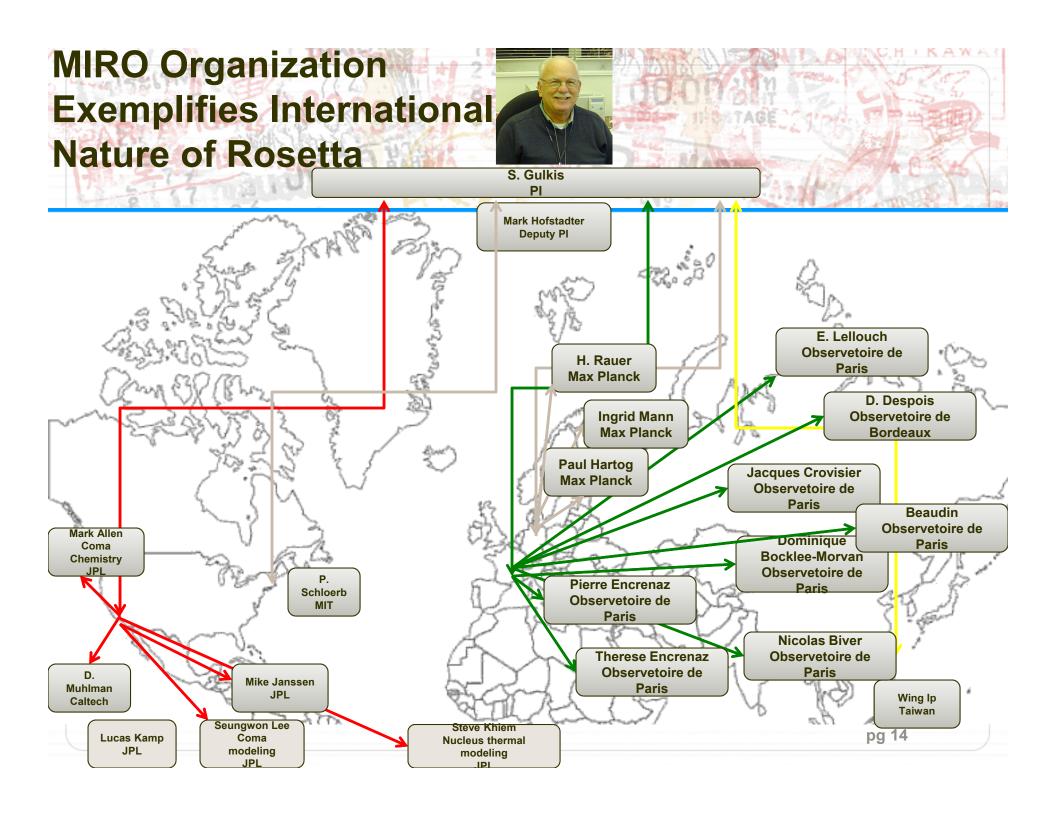
Science Support

VIRTIS
Robert Carlson, Co-I (JPL)
Michael Combi, Co-I (Univ of Michigan)
Uwe Fink, Co-I (Univ of Arizona)
Tom McCord, Co-I (Bern Flight Center)
Radio Science
John Anderson, Co-I (SwRI)
Essam Marouf, Co-I (San Jose State Univ.)
OSIRIS

Mike A'Hearn, Co-I (Univ. of Maryland)







NASA International Mission Management Lessons

ITAR

- Important to write a good LOA up-front and cover future issues
- Staff must be well trained in export controls
- Access to knowledgeable advisors
- Find ways of sharing information

• If you are a junior partner, behave like one

- Always give credit to the senior partner
- Do not try to impose your processes, methods and bureaucracy
- Don't be a burden with knowledge, reporting, reviews
- If you are a senior partner use junior's expertise
- DSN gives new capability to any partner





NASA International Mission Management Lessons

- ESA's has more decentralized management authority than NASA
 - Mission funding comes from different countries
 - Mission functionality is divided among many countries
 - Ground segment Madrid
 - Operations Darmstadt
 - Project Scientist Nordwijk
 - Antenna Madrid
 - Instruments Germany, Italy, Switzerland, Great Britain
 - Archiving Paris
- Get used to lots, and lots of emails, no phone
- Get used to meetings that start at midnight
- Develop relationships and trust







- Flexible PI and Co-I selection process needed for long duration missions
 - Over a decade scientists' interests change
 - New superstars appear on the research horizon
 - Need an easy system for updating the science team membership
- Funds don't cross the water
 - Must have budget reserves that allow flexibility with foreign PI's
 - Continuous effort to align the US scientists with their ESA PI's
- Great benefit to NASA from access to 100% Rosetta data despite 10% budget contribution





NASA decided to participate in the Rosetta mission a decade ago, but it fulfills recommendations for the next decade:

- Low cost (to NASA)
- International cooperation
- Visiting a small primitive body
- Exciting, challenging objectives